

QG2

SPINFOAM COSMOLOGY

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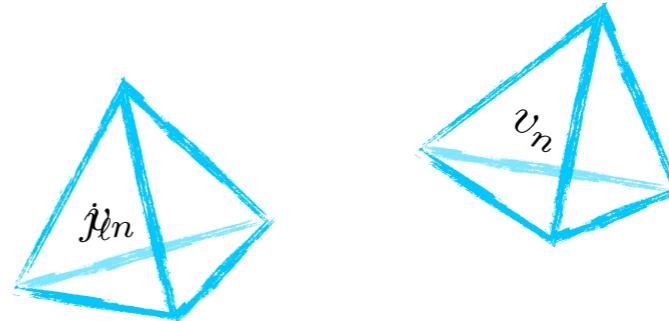
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PLAN OF THE TALK

- **SPINFOAM** *I present the details of the calculation for a generic regular graph.*
- **COSMOLOGY** *I discuss the simplified framework provided by cosmology.*
- **OPEN ISSUES** *New numerics validate the previous results, but also demand us to rethink about some steps taken.*
- **ACHTUNG** *technical talk! for a general presentation, see my talk at QG4 session on Friday*

COHERENT STATES

- Spinnetwork states $|\Gamma, j_\ell, v_n\rangle$



Bianchi, Magliaro, Perini

- Coherent states $|\Gamma, z_\ell, \vec{n}_\ell, \vec{n}'_\ell\rangle$

- Geometrical interpretation for the labels $(z_\ell, \vec{n}_\ell, \vec{n}'_\ell)$:

Freidel, Speziale

$\vec{n}_\ell, \vec{n}'_\ell$ are the 3d normals to the faces of the cellular decomposition;

$Im(z_\ell) \leftrightarrow$ curvature at the faces and $Re(z_\ell) \leftrightarrow$ area of the face

$$Re(z_\ell) = \theta(\gamma K + \Gamma)$$

- Hom&Iso coherent states $|\Gamma, z\rangle$

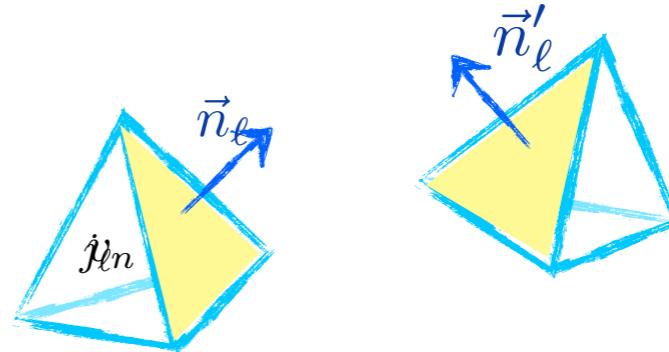
Marcianò, Magliaro, Perini, Rovelli, FV

$\vec{n}_\ell, \vec{n}'_\ell$ fixed by requiring a regular cellular decomposition

- in terms of the scale factor $Re(z) \sim \dot{a}$ and $\sqrt{Im(z)} \sim a$

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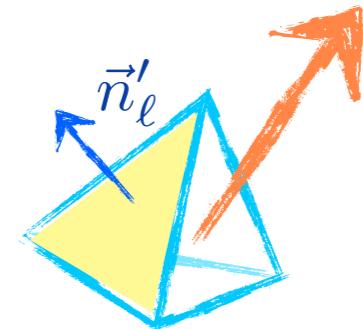
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VERTEX AMPLITUDE

$$\langle (j, \gamma j); j', m' | Y | j, m \rangle = \delta_{p, \gamma j} \delta_{kj} \delta_{jj'} \delta_{mm'}$$

$$W_v(h_\ell) = \int_{SL(2, \mathbb{C})} \prod_{n=1}^{N-1} dG_n \prod_{\ell=1}^L P(h_\ell, G_\ell)$$

kernel of the map:

$$\begin{aligned} Y : \quad \mathcal{H}^{(j)} &\longrightarrow \mathcal{H}^{(j, \gamma j)} \\ |j, m\rangle &\qquad |(j, \gamma j); j, m\rangle \end{aligned}$$

$$P(h_\ell, G_\ell) = \sum_{j_\ell} (2j_\ell + 1) D^{(j_\ell)}(h_\ell)_m^{m'} D^{(\gamma j_\ell, j_\ell)}(G_\ell)_{jm'}^{jm}$$

• coherent states

$$\psi_{H_\ell}(h_\ell) = \int_{SU(2)^N} dg_n \prod_{\ell=1}^L \sum_{j_\ell} (2j_\ell + 1) e^{-2t\hbar j_\ell(j_\ell + 1)} \text{Tr} [D^{(j_\ell)}(h_\ell)]$$

$$H_\ell \in SL(2, \mathbb{C})$$

$$P_t(H_\ell, G_\ell) = \int dh_\ell K_t(h_\ell, H_\ell) P(h_\ell, G_\ell)$$

$$= \sum_{j_\ell} (2j_\ell + 1) e^{-2t\hbar j_\ell(j_\ell + 1)} \text{Tr} [D^{(j_\ell)}(H_\ell) Y^\dagger D^{(\gamma j_\ell, j_\ell)}(G_\ell) Y]$$

SADDLE POINT

- Coherent states introduces

$$D^{(j)}(H_\ell(z)) = D^{(j)}(R_{\vec{n}_s}) \ D^{(j)}(e^{-iz\frac{\sigma_3}{2}}) \ D^{(j)}(R_{\vec{n}_n}^{-1})$$

$$D^{(j)}(e^{-iz\frac{\sigma_3}{2}}) = \sum_m e^{-izm} \ |m\rangle\langle m|$$

Im(z) large: $D^{(j)}(e^{-iz\frac{\sigma_3}{2}}) \approx e^{izj} \ |j\rangle\langle j|$

$$D^{(j)}(H_\ell(z)) = D^{(j)}(R_{\vec{n}_s}) \ e^{-izj} |j, +j\rangle\langle j, +j| D^{(j)}(R_{\vec{n}_n}^{-1}) = e^{-izj} |j, \vec{n}_\ell\rangle\langle j, \vec{n}_\ell|$$

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- $\text{Tr} [D^{(j_\ell)}(H_\ell) Y^\dagger D^{(\gamma j_\ell, j_\ell)}(G_\ell) Y]$

$$e^{-izj} \langle j, \vec{n}_\ell | \ Y^\dagger D^{(\gamma j_\ell, j_\ell)}(G_\ell) Y | j, \vec{n}_\ell \rangle$$

$$\int_{SL(2, \mathbb{C})} \prod_{n=1}^{N-1} dG_n \sum_{j_\ell} \prod_{\ell=1}^L (2j_\ell + 1) \ e^{-2t\hbar j_\ell(j_\ell+1)} \ e^{-izj} \langle (\gamma j, j); j, \vec{n}_\ell | \ D^{(\gamma j_\ell, j_\ell)}(G_\ell) | (\gamma j, j); j, \vec{n}_\ell \rangle$$

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$$\sum_{j_\ell} \prod_{\ell=1}^L (2j_\ell + 1) \ e^{-2t\hbar j_\ell(j_\ell+1)} e^{-izj} \int_{SL(2,\mathbb{C})} \prod_{n=1}^{N-1} dG_n \prod_{\ell=1}^L \langle (\gamma j, j); j, \vec{n}_\ell | \ D^{(\gamma j_\ell, j_\ell)}(G_\ell) | (\gamma j, j); j, \vec{n}_\ell \rangle$$

- $j_\ell \rightarrow \alpha j_\ell$ and $\alpha \gg 1$

$$\Omega(j_\ell) \approx \frac{1}{\alpha^3 \sqrt{\det \text{Hess}(j_\ell)}} \ e^{-\frac{1}{2} i j_\ell \theta}$$

EVALUATION OF THE AMPLITUDE

$$W(z) = \sum_{j_\ell} \prod_{\ell=1}^L \frac{1}{\alpha^3 \sqrt{\det \text{Hess}(j_\ell)}} (2j_\ell + 1) e^{-2t\hbar j_\ell(j_\ell+1)-izj_\ell} e^{-\frac{1}{2}ij_\ell\theta}$$

$$\theta(\gamma K + 1) - \theta = 0$$

- Gaussian sum peaked at j_o for all j_ℓ $j \sim j_o + \delta j$
 - max (real part of the exponent) gives where the gaussian is peaked; $j_o \sim \text{Im } \tilde{z}/4\text{th}$
 - imaginary part of the exponent = $2k\pi$ gives where the gaussian is not suppressed. $\text{Re } \tilde{z} = 0$ $\dot{a} \sim 0$
 - We obtain Minkowski space!

$$W(z) = \left(\sqrt{\frac{\pi}{t}} e^{-\frac{\tilde{z}^2}{8t\hbar}} 2j_o \right)^L \frac{N_\Gamma}{j_o^3}$$

EVALUATION OF THE AMPLITUDE

- Cosmological constant

$$W(z) = \sum_j (2j+1) \frac{N_\Gamma}{j^3} e^{-2t\hbar j(j+1)-izj-i\lambda v_o j^{\frac{3}{2}}}$$

intertwiner $v_e \sim v_o j^{3/2}$

$$i\lambda v_o j^{\frac{3}{2}} \sim i\lambda v_o j_o^{\frac{3}{2}} + \frac{3}{2} i\lambda v_o j_o^{\frac{1}{2}} \delta j$$

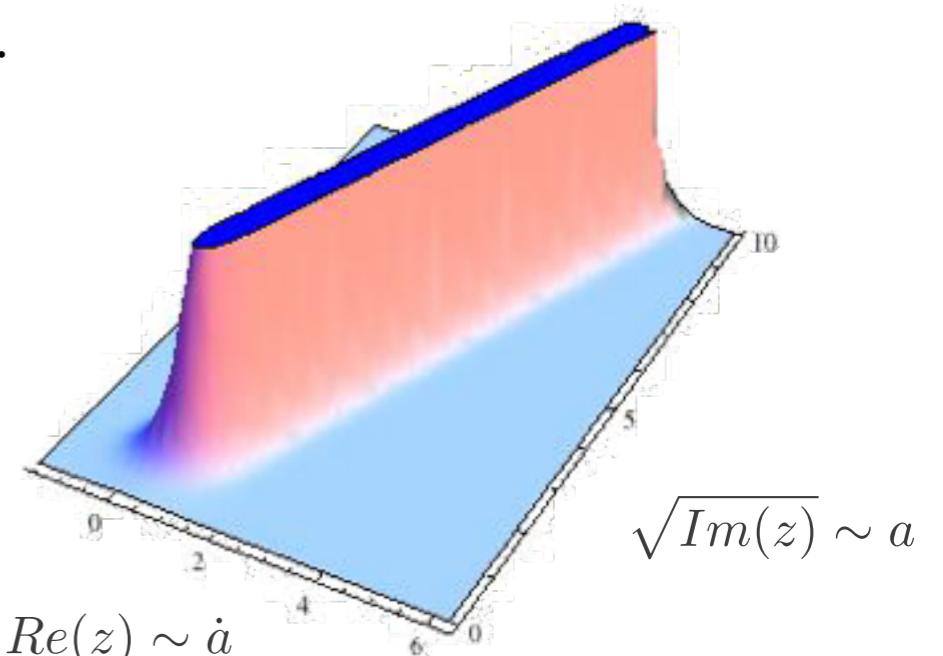
- the gaussian is peaked on

$$j_o = \frac{Im(z)}{4t\hbar}$$

- the gaussian is not suppressed for $Re(z) + \lambda v_o j^{\frac{1}{2}} = 0$.

$$\frac{Re(z)^2}{Im(z)} = \frac{\lambda^2 v_o^2}{4t\hbar} \rightarrow \left(\frac{\dot{a}}{a}\right)^2 = \frac{\Lambda}{3}$$

$$\Lambda = \text{const } \lambda^2 G^2 \hbar^2$$



RESULTS and OPEN ISSUES

- *There are approximations in the quantum theory that yield cosmology.*
- *Coherent states for cosmology.*
- *There is a simple way to add the cosmological constant to LQG dynamics.*
- *The theory recover general relativity in the semiclassical limit, also for non-trivial solutions.*
- *Connecting canonical and covariant in loop cosmology.*
- *We can couple fermions in the full theory.*
- *Are these approximations viable?*
- *Is there any relation between coherent states in a truncation and embedding?*
- *There is a complicated way to add the cosmological constant using q -deformed groups.*
- *We want to go beyond semiclassicality, studying quantum correction.*
- *Derivation of $\bar{\mu}$ -scheme in the covariant theory.*
- *Matter in Spinfoam Cosmology?*